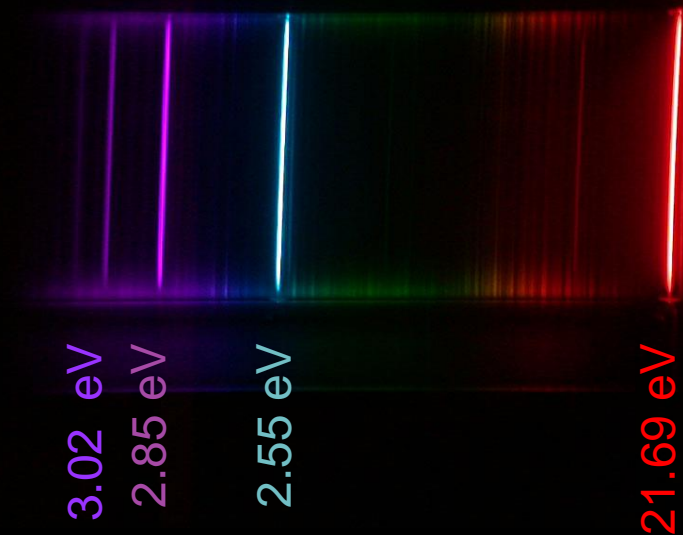


Wed.	8.1-.3, (.8,.9) Photons & Quantization Quiz 7	RE 8.a
Lab	L7 Microscopic Energy Transfer	
Fri.	8.4-.7 More Energy Quantization	RE 8.b
Mon.	9.1-.2, (.8) Momentum and Energy in Multi-	RE 9.a
Tues.	particle Systems	HW8: Ch 8 Pr's 21, 23, 27(a-c)
Wed.	9.3 Rotational Energy Quiz 8	RE 9.b
Lab	Review Exam 2 (Ch 5-8)	
Fri.	Exam 2 (Ch 5-8)	Practice Exam 2 (bring to class)

*Shedding light on atomic energy levels
(segment of Hydrogen spectrum)*



Where we've been:

Energy on the macroscopic scale

$$\Delta K_{paper} + \Delta U_{p\&E}$$

$$\Delta\left(\frac{1}{2}m_p v_p^2\right) + \Delta(mgh)$$

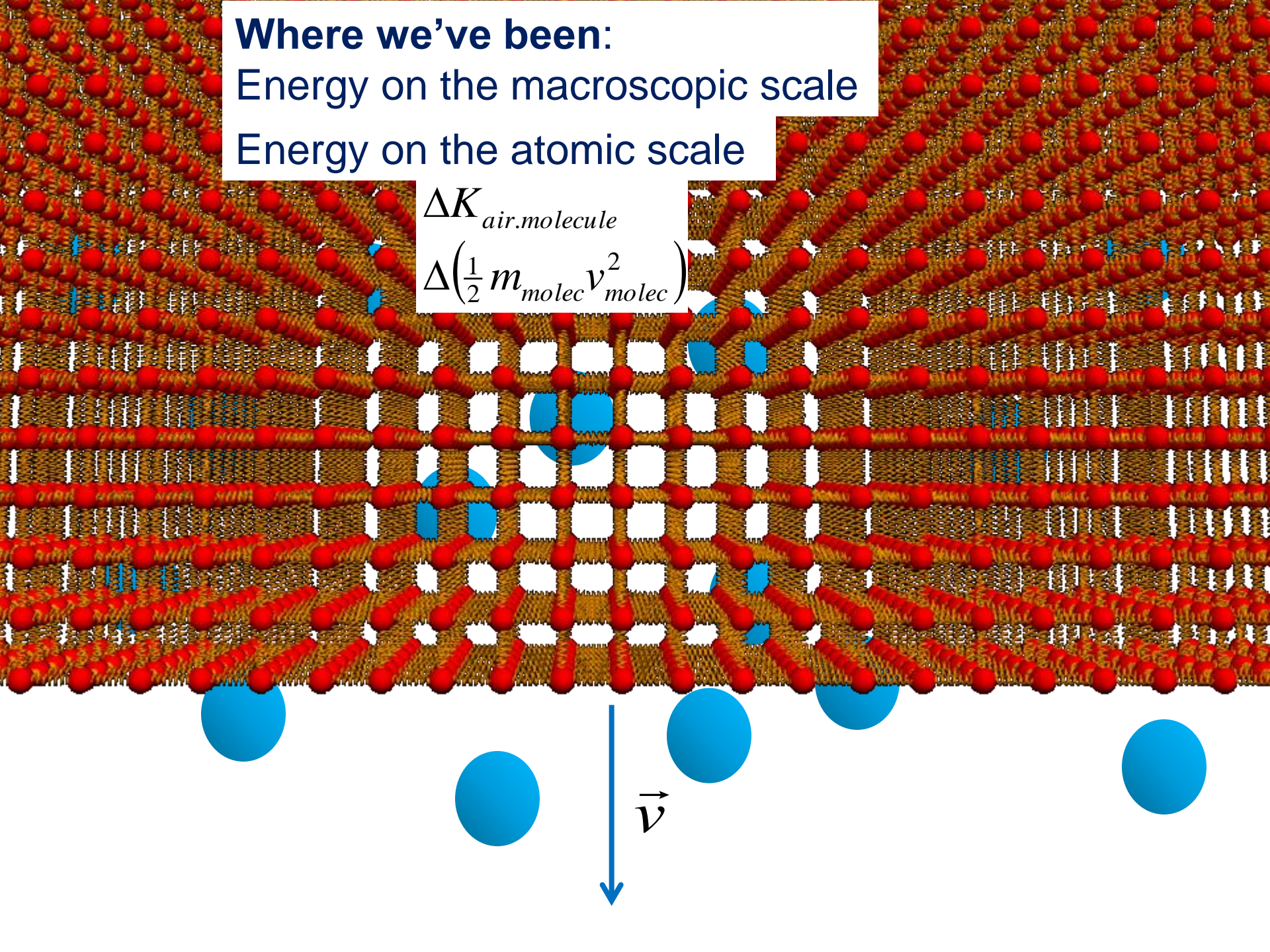
Where we've been:

Energy on the macroscopic scale

Energy on the atomic scale

$$\Delta K_{air.molecule}$$

$$\Delta \left(\frac{1}{2} m_{molec} v_{molec}^2 \right)$$



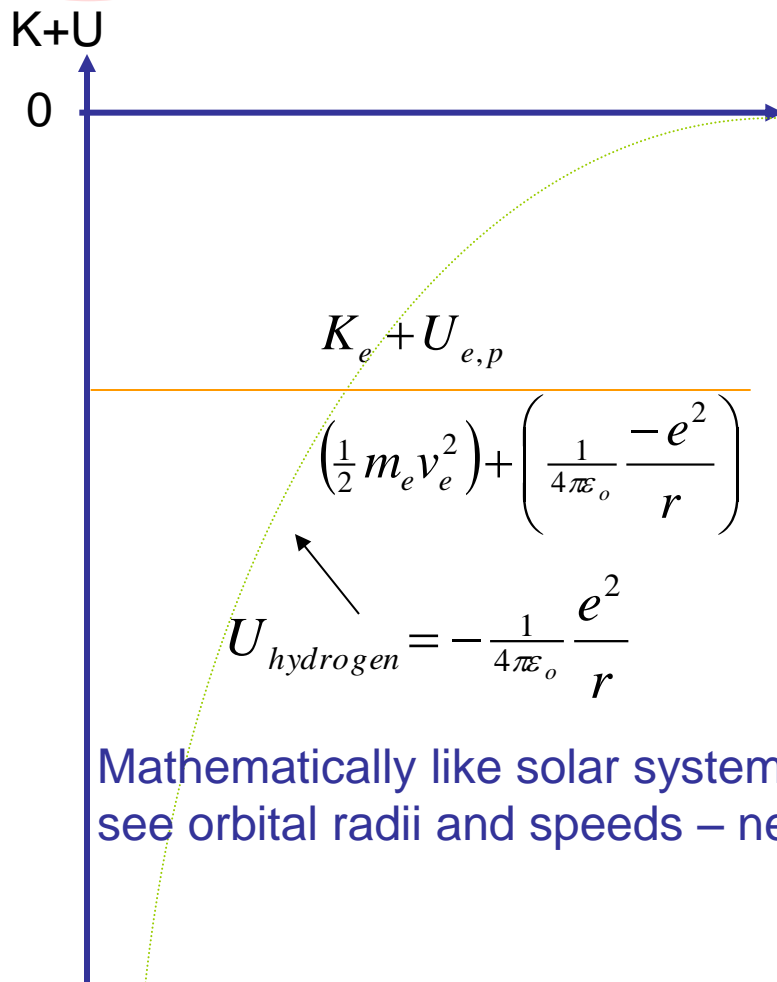
Where we've been:

Energy on the macroscopic scale

Energy on the atomic scale

Where we're going:

Energy on the electronic scale

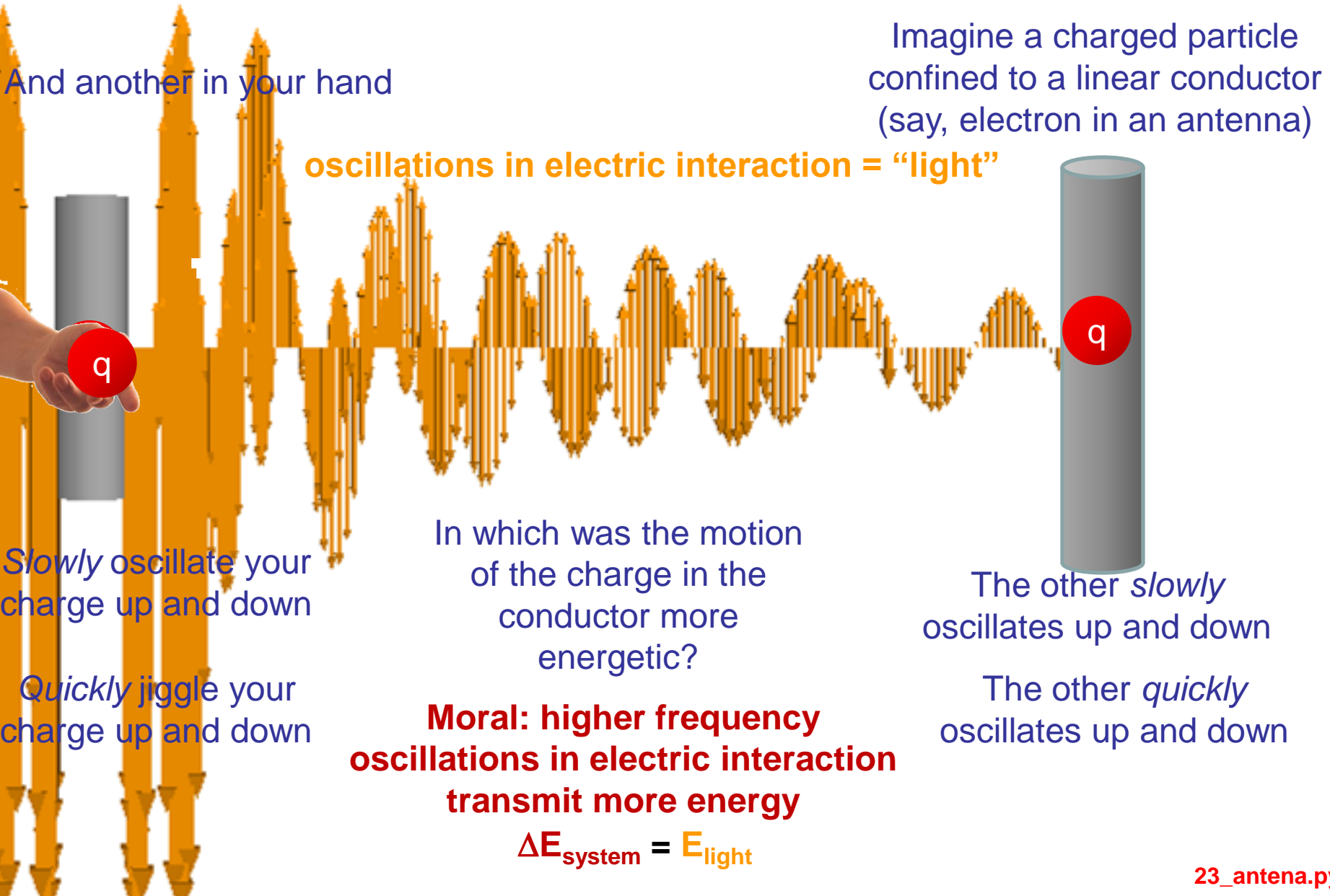


$$E_H \approx m_e c^2 + m_p c^2 + (K_e + U_{e,p})$$

Mathematically like solar system, but much too small and delicate to directly see orbital radii and speeds – need another way to deduce

Spectroscopy

Light and energy

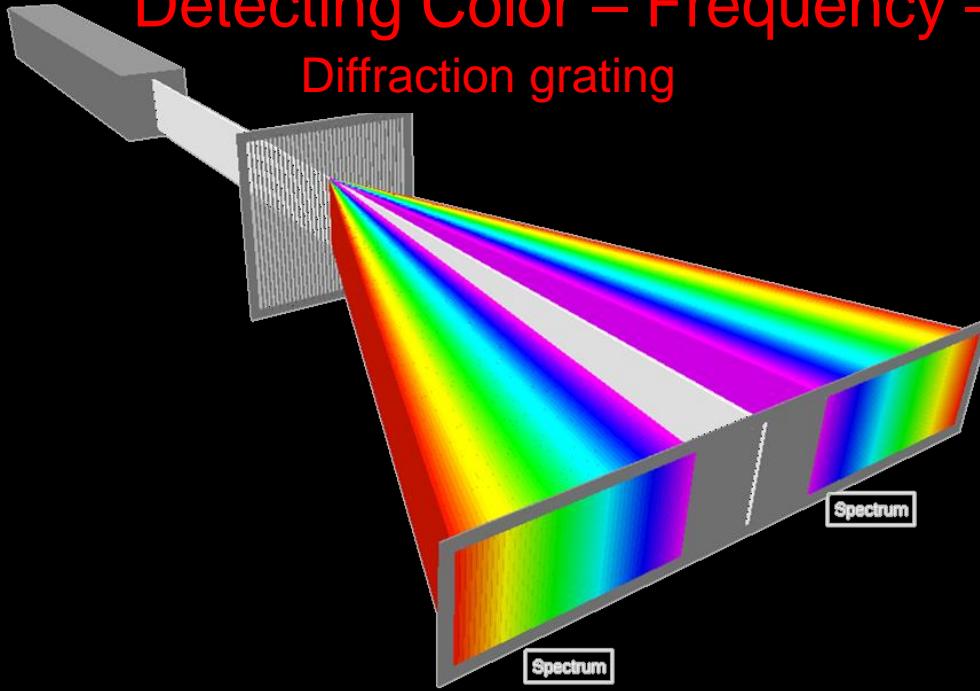


White light
source

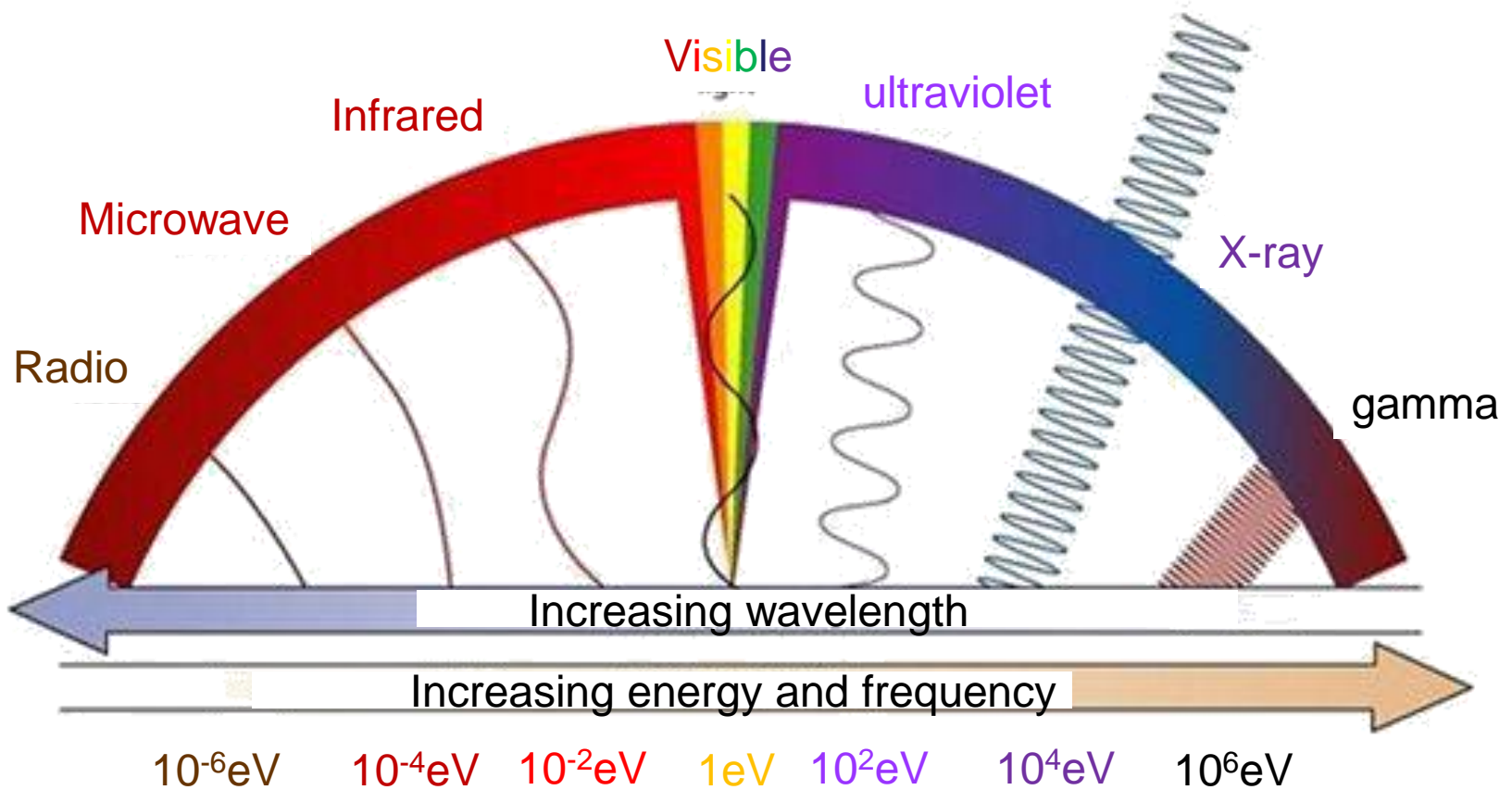
Spectrometer:

Detecting Color – Frequency – Energy

Diffraction grating



Spectrometer: Detecting Color – Frequency – Energy



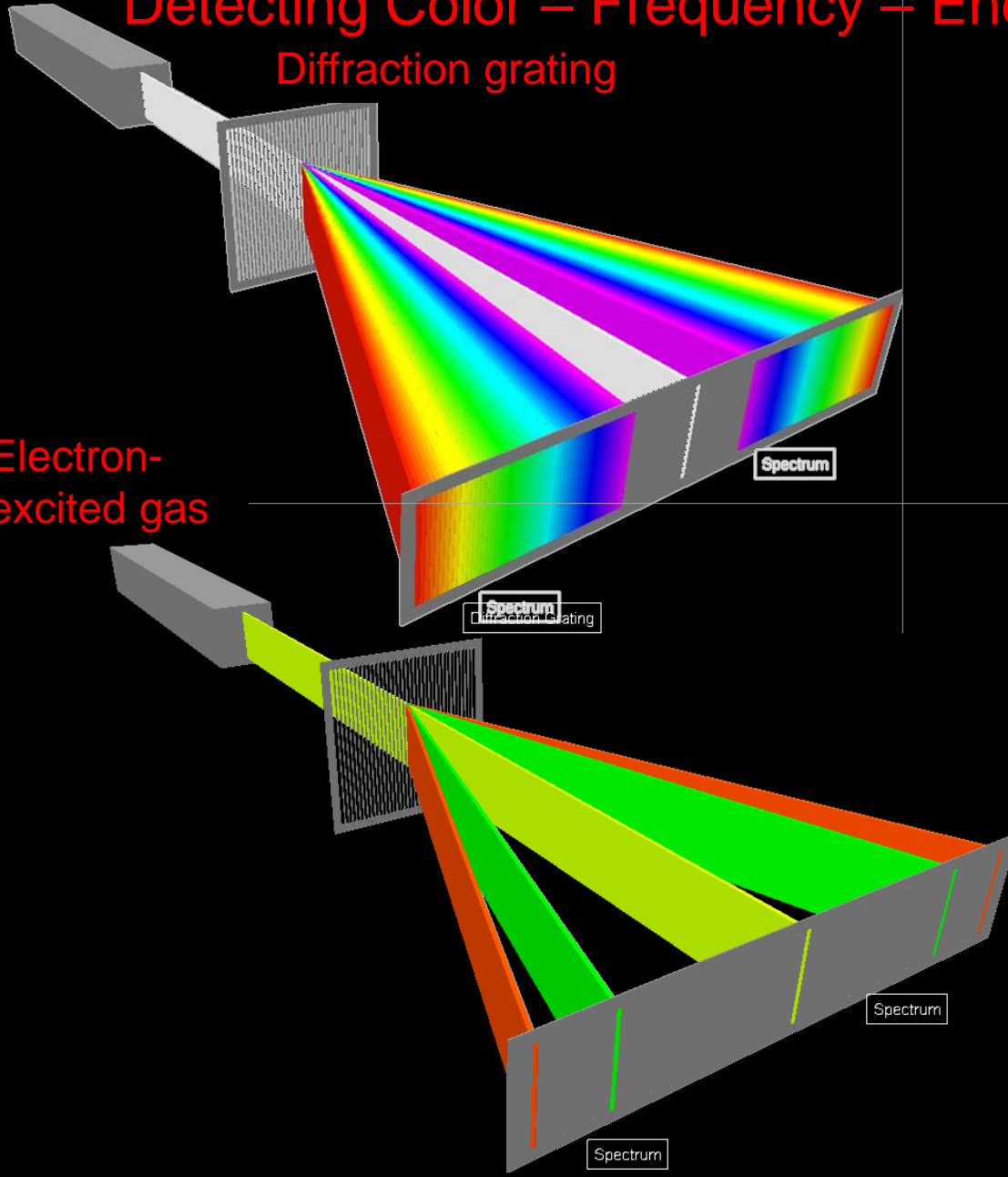
White light source

Spectrometer:

Detecting Color – Frequency – Energy

Diffraction grating

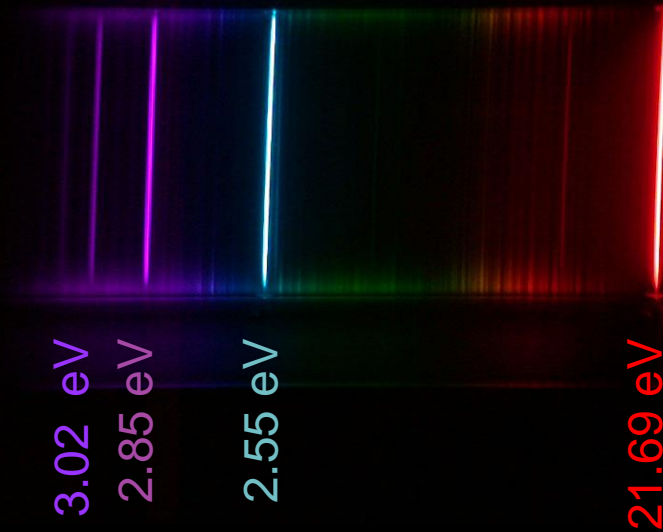
Electron-excited gas



Spectrometer: Detecting Color – Frequency – Energy Hydrogen

$$E_{light.from.H} = 13.6eV \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

Where $n_2 > n_1$, integers



Only specific E_{light} , so...

Only specific ΔE_H , so...

Only specific E_H

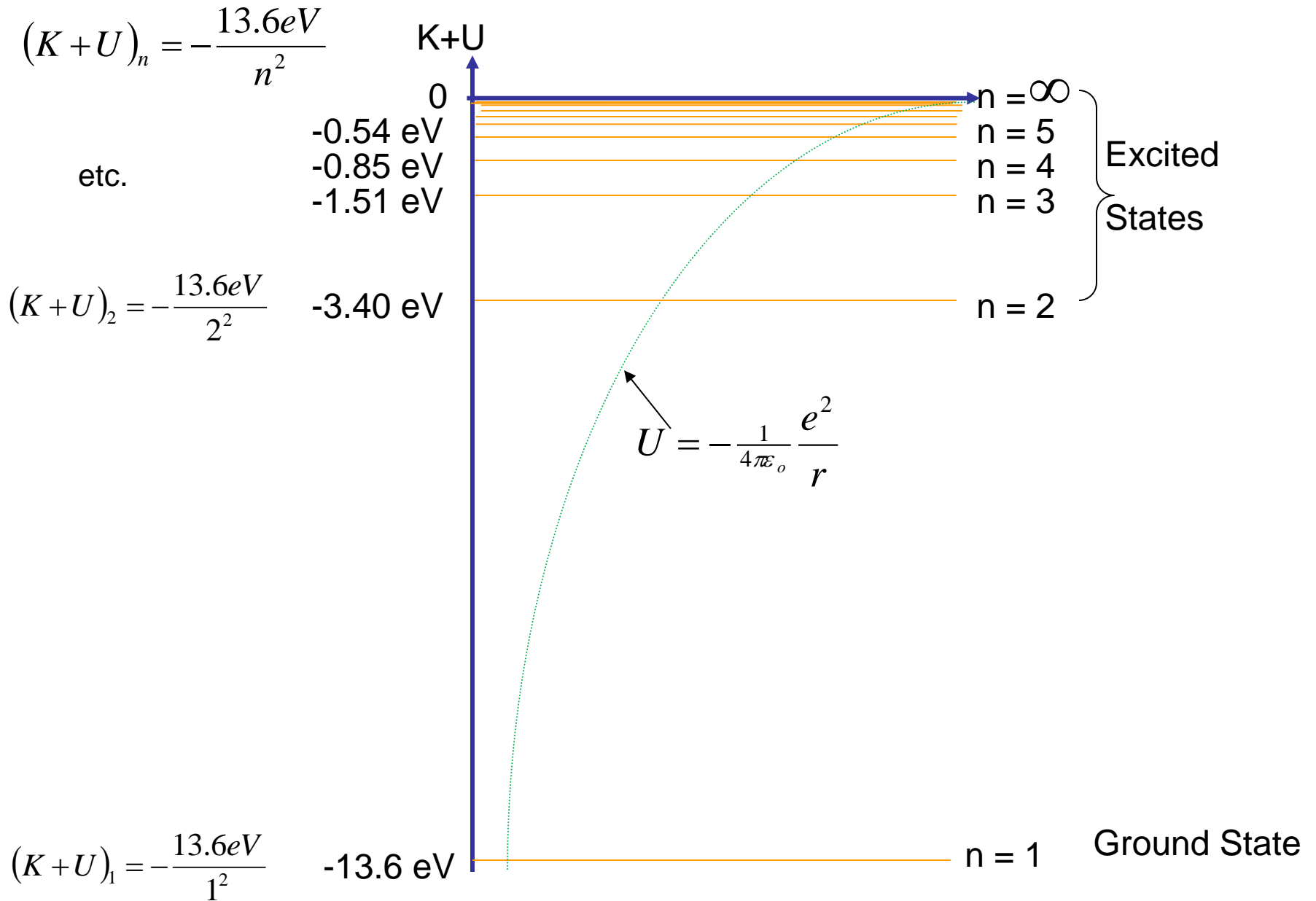
$$E_{H,n} \approx m_e c^2 + m_p c^2 + (K_e + U_{e,p}) = m_e c^2 + m_p c^2 + \left(\frac{-13.6eV}{n^2} \right)$$

$$E_{H,n} \approx m_e c^2 + m_p c^2 + (K_e + U_{e,p}) = m_e c^2 + m_p c^2 + \left(\frac{-13.6 \text{ eV}}{n^2} \right)$$

Q8.2.a: What is (K+U) of the sixth electronic energy level (n=6) of a hydrogen atom?

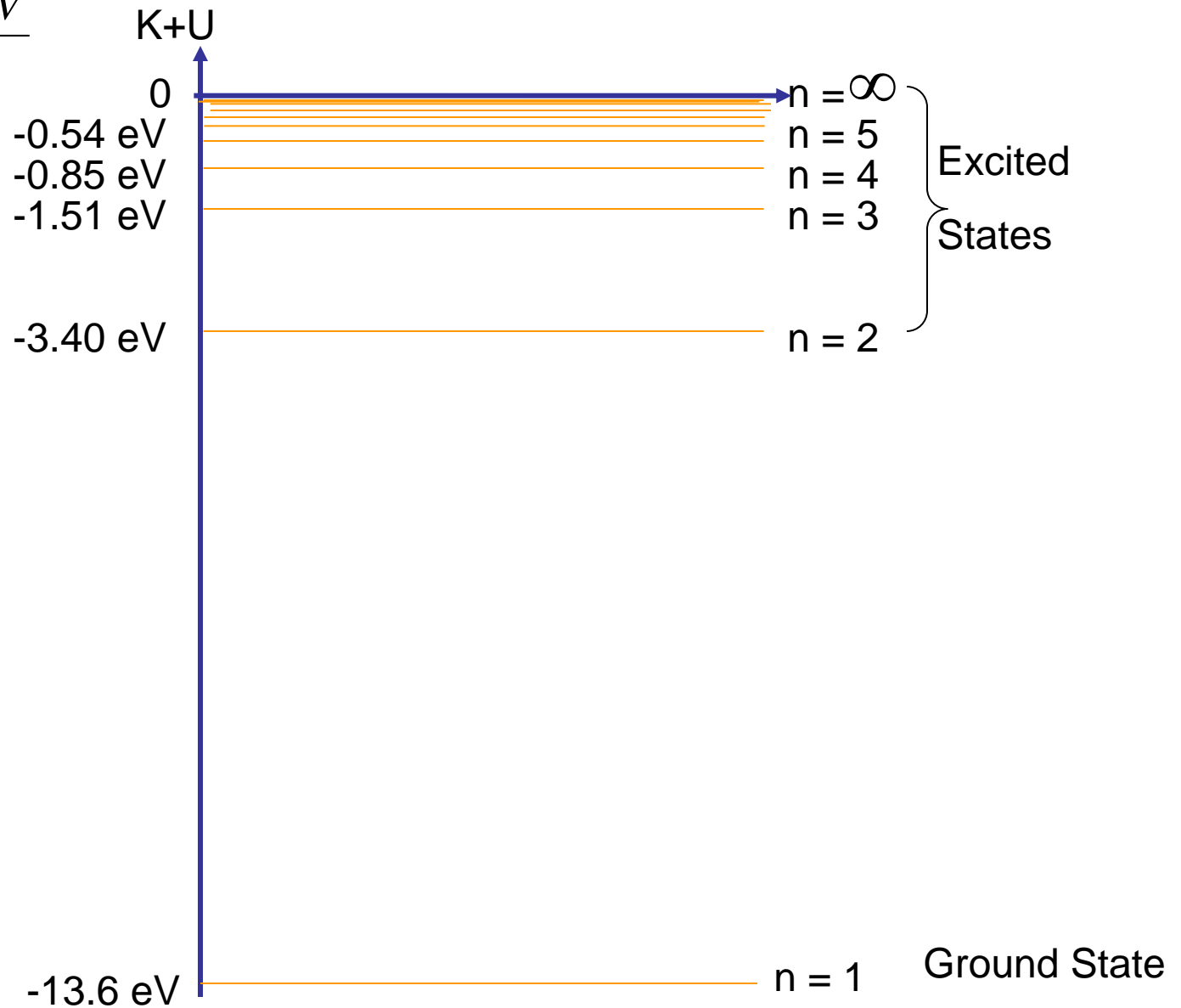
- 1) - 13.6 eV**
- 2) - 2.27 eV**
- 3) + 2.27 eV**
- 4) - 0.38 eV**
- 5) + 0.38 eV**

Hydrogen Energy Levels



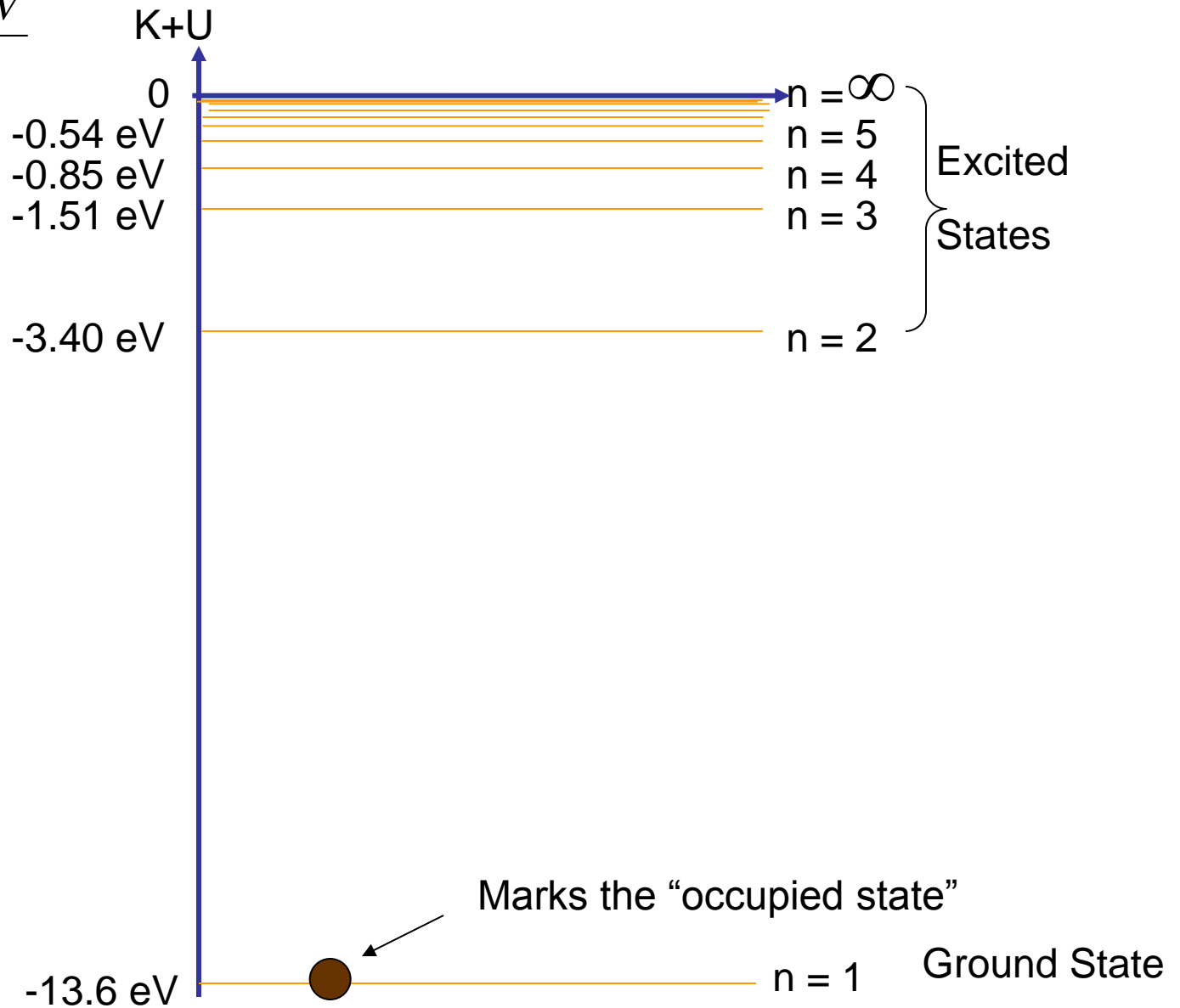
Hydrogen Energy Levels: Excitation

$$(K + U)_n = -\frac{13.6\text{eV}}{n^2}$$



Hydrogen Excitation: 1st in ground state

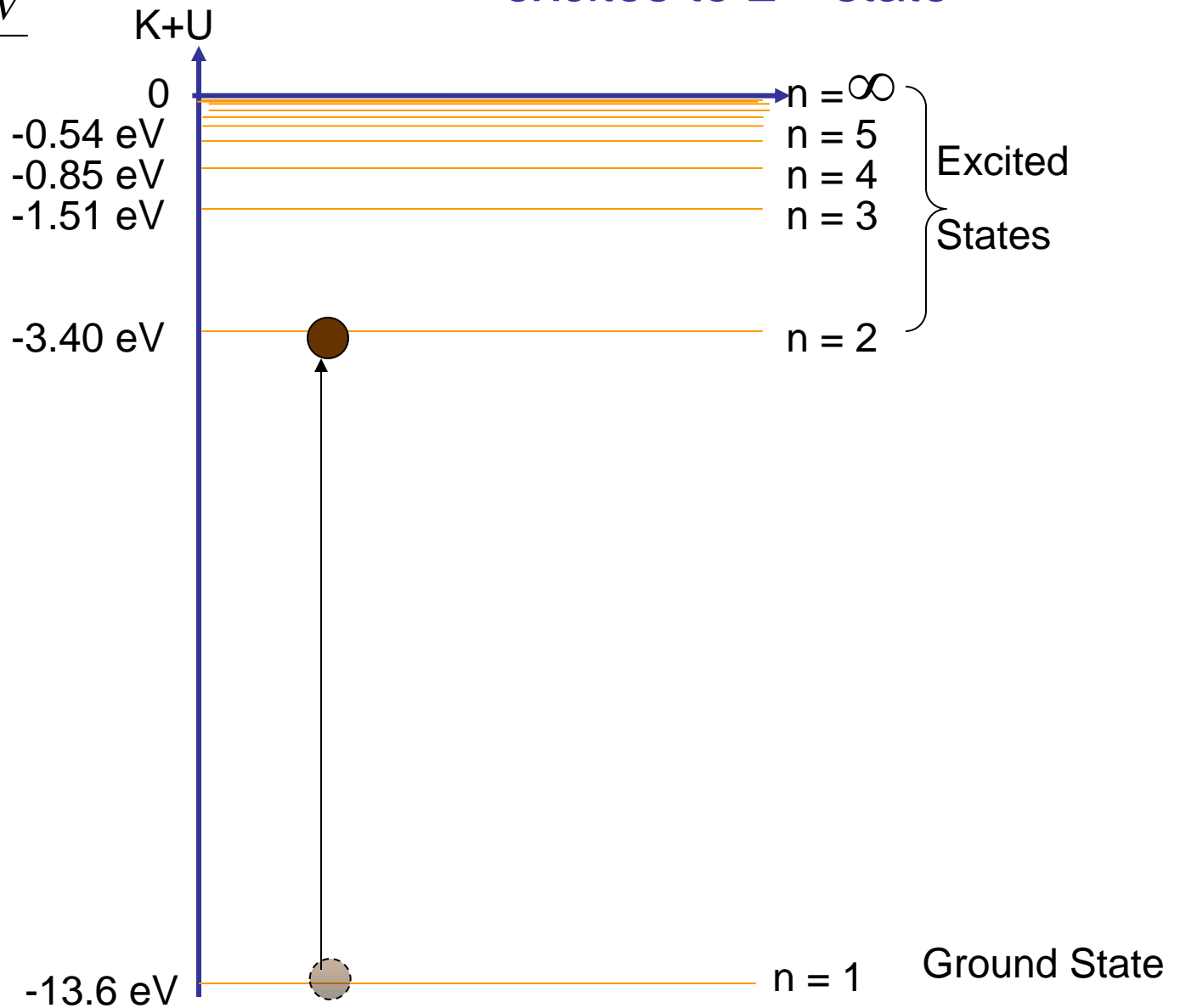
$$(K + U)_n = -\frac{13.6\text{eV}}{n^2}$$



Hydrogen Excitation: 2nd Adsorbs energy from Collision

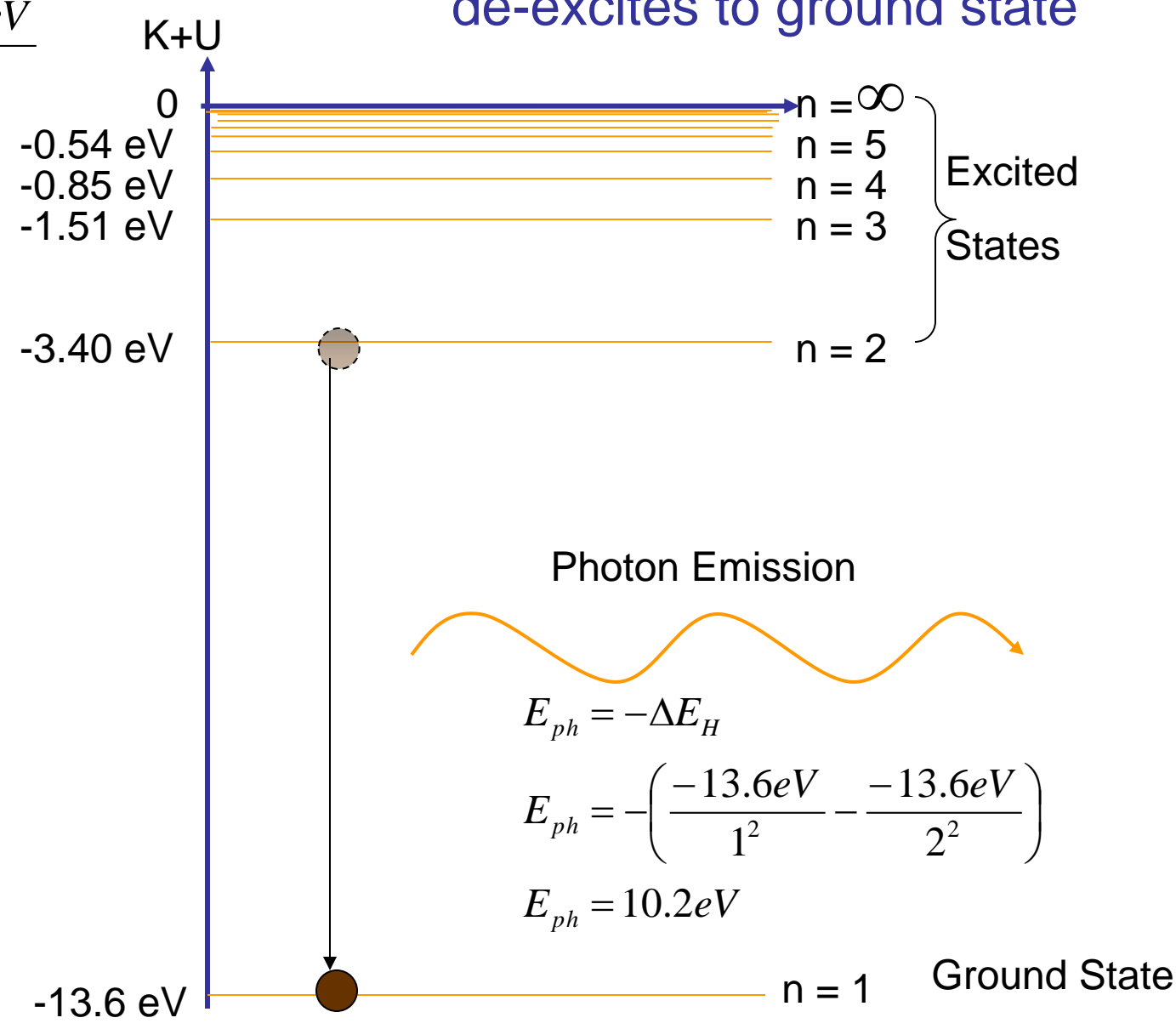
excites to 2nd state

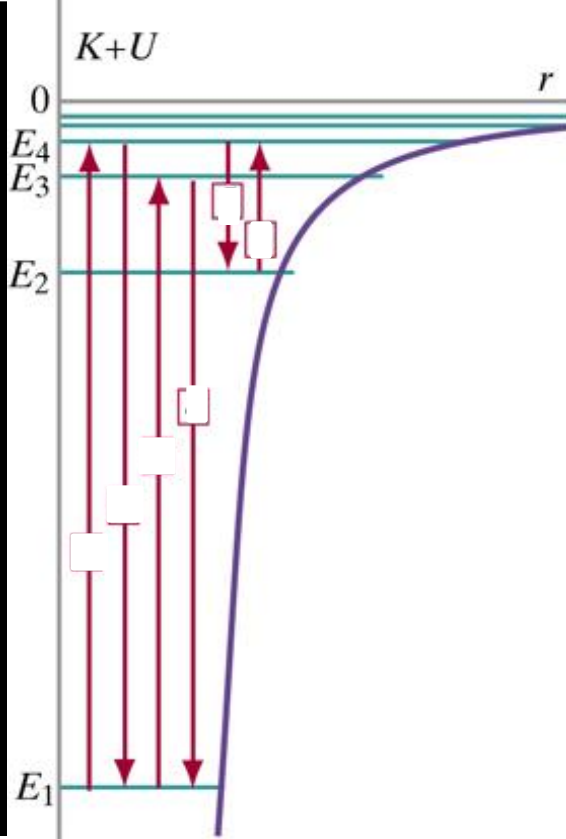
$$(K + U)_n = -\frac{13.6\text{eV}}{n^2}$$



Hydrogen Excitation: 3rd Loses Energy by photon emission, de-excites to ground state

$$(K + U)_n = -\frac{13.6\text{eV}}{n^2}$$





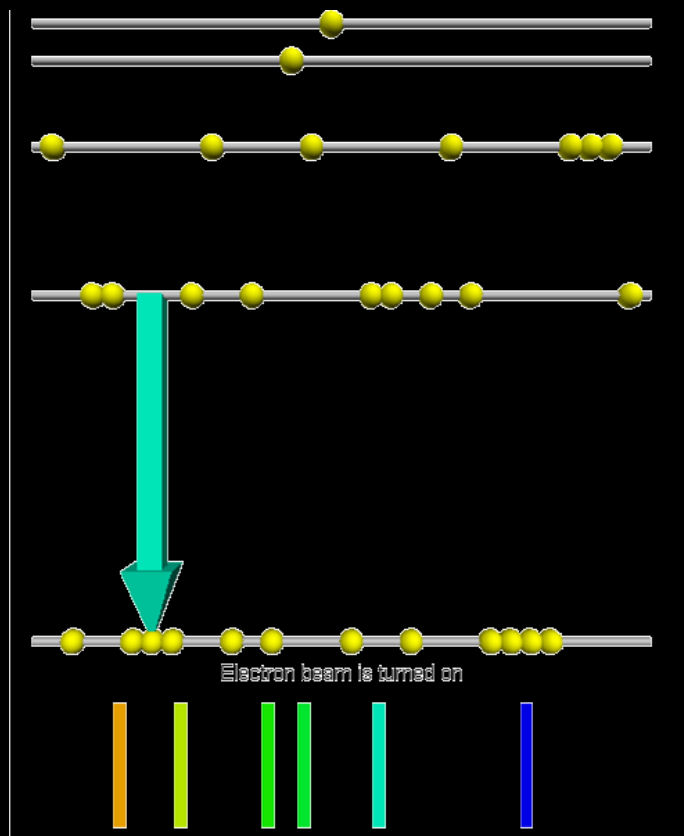
Q.8.2.c

A hydrogen atom is initially in the excited state (N = 4). It emits a photon and ends up in the state (N = 2).

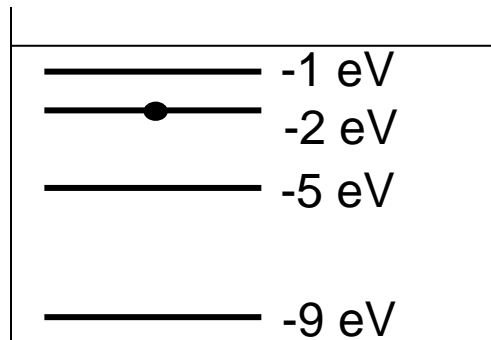
What is the energy of the emitted photon?

- 1) $E_4 - E_2$
- 2) $E_2 - E_4$
- 3) E_4
- 4) E_2
- 5) $|E_4|$
- 6) $|E_2|$

Example Atoms in Gas-Discharge Tube



Emission Example Here are the quantized energy levels (K+U) for some atomic or molecular object. If the object is excited to the third level (marked by a dot), what are the possible energies of photons that may be emitted?



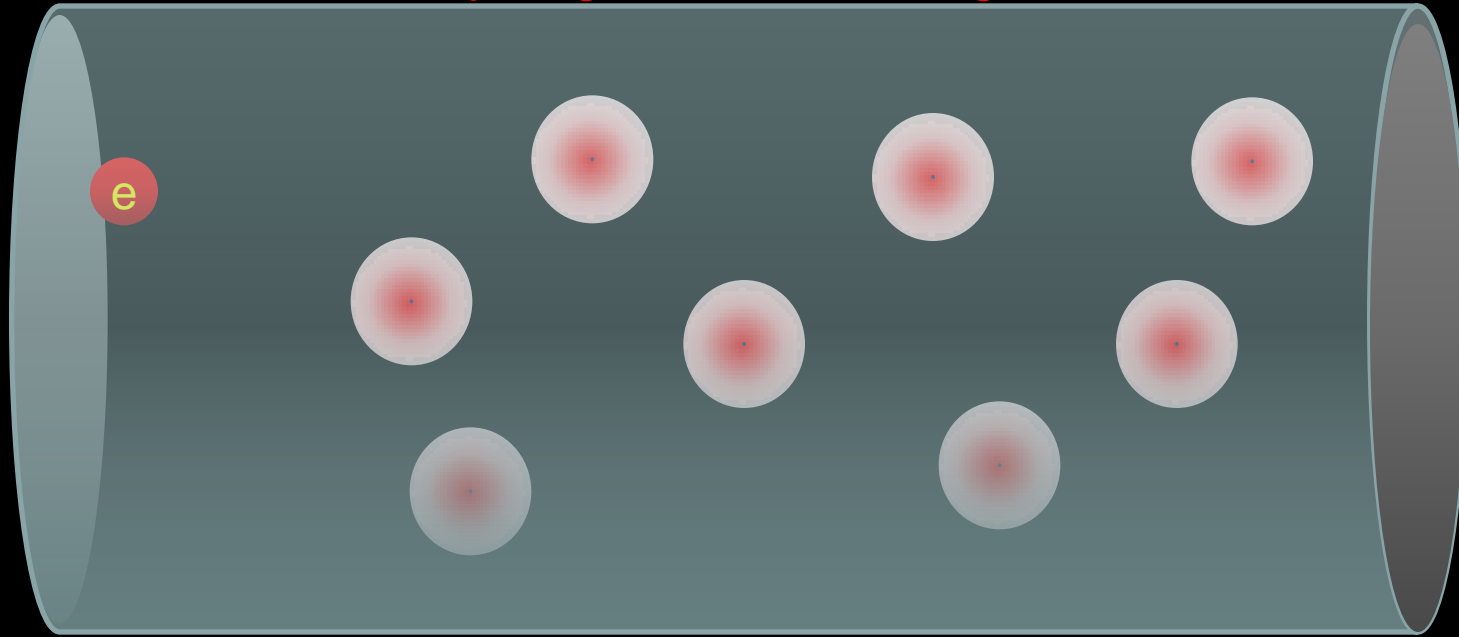
Say we have a very hot gas of such atoms, so the electron can get knocked into any of the levels. How many distinct energies can emitted photons have? (note: you'll want to actually determine each value to check for duplicates).

- a) 2
- b) 3
- c) 4
- d) 5
- e) 6
- f) 7
- g) 8

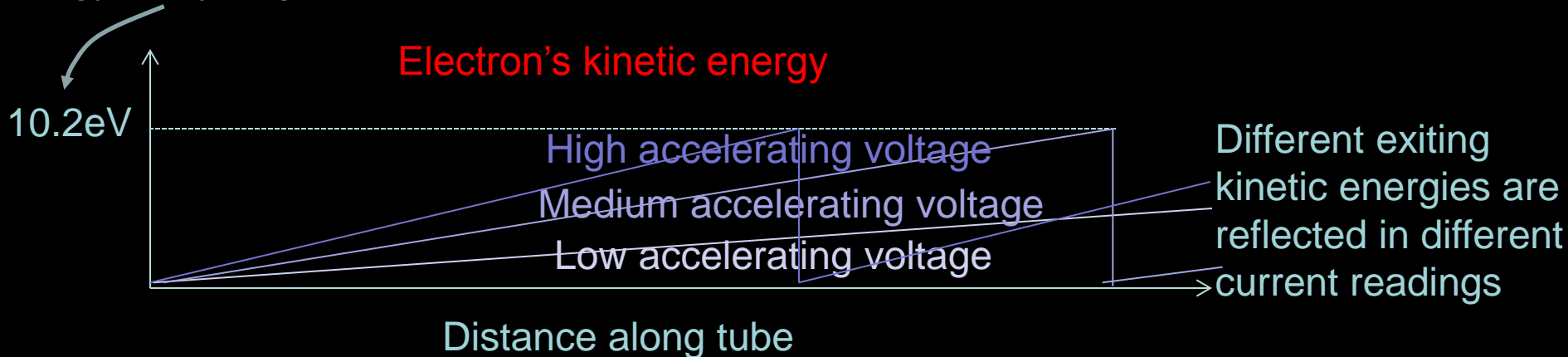
Frank-Hertz Experiment

Monitoring electron beam's *loss* of energy to the atoms

Hydrogen Gas Discharge Tube

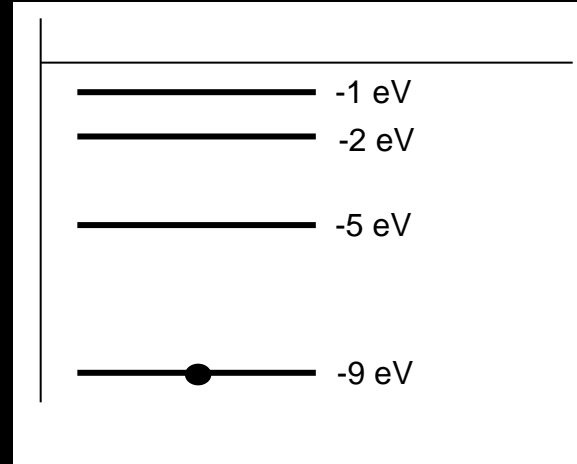


Energy of Hydrogen's $n=1$ to 2 transition

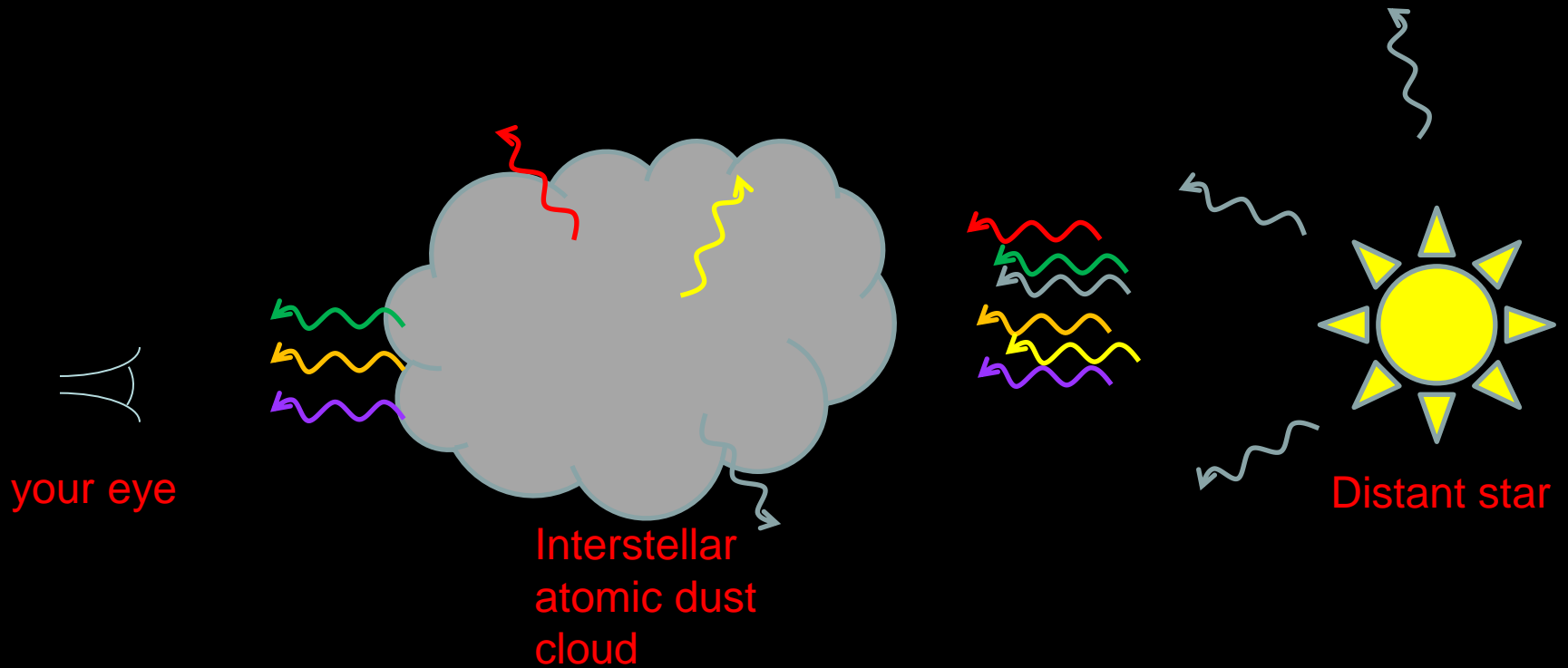


Here are the quantized energy levels (K+U) for an atomic or molecular object, and the object is in the "ground state" (marked by a dot). An electron with kinetic energy 6 eV is fired at the object and excites the object to the -5 eV energy state. What is the remaining kinetic energy of this electron?

- a) 9 eV
- b) 6 eV
- c) 4 eV
- d) 3 eV
- e) 2 eV



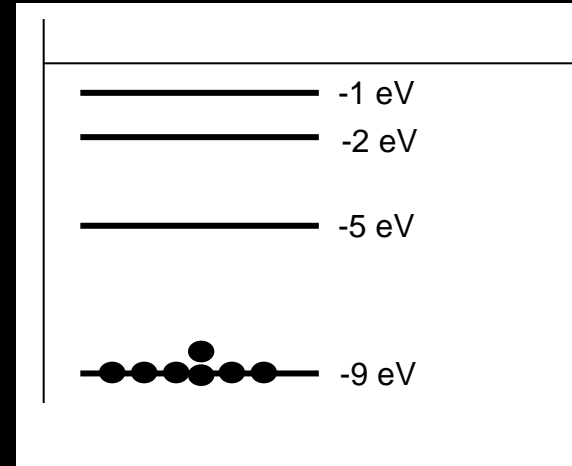
Absorption Spectrum



Colors / energies of light from the star that interact with cloud's atoms scatter; it's depleted from the star light you see.

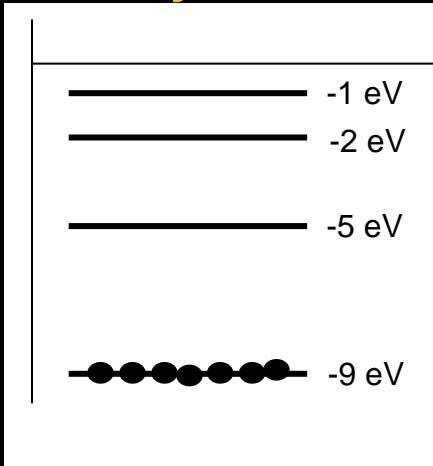
A collection of some atoms objects is kept **very cold**, so that all the objects are in the ground state. Light consisting of photons with a range of energies from 1 to 7.5 eV passes through this collection of objects. What photon energies will be depleted from the light beam (“dark lines”)?

- a) 2 eV, 5 eV, 9 eV
- b) 3 eV, 4 eV
- c) 0.5 eV, 3 eV, 4 eV
- d) 4 eV, 7 eV
- e) 3 eV, 4 eV, 7 eV



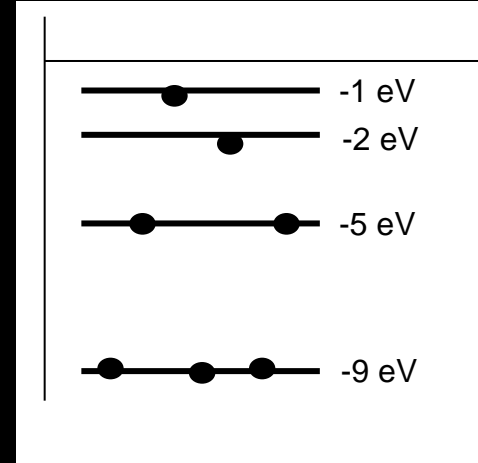
Temperature Effects on Absorption Spectrum

T very low



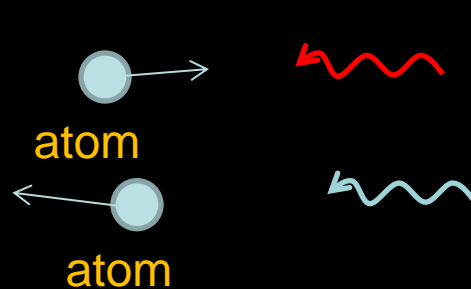
All atoms initially in ground state; only absorption lines for transitions from it

T very high



Many states have some atoms; you see absorption lines between many states

T medium



More energetic collision; lower-energy (redder) light excites transition

Less energetic collision; higher (bluer) light excites transition

Thermally broadened Spectral lines

Wed. Lab Fri.	8.1-.3, (.8,.9) Photons & Quantization Quiz 7 L7 Microscopic Energy Transfer 8.4-.7 More Energy Quantization	RE 8.a RE 8.b
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*Shedding light on atomic energy levels
(segment of Hydrogen spectrum)*

